

## DISOCIACION COULOMBIANA DEL $^{27}\text{P}$

S.Beceiro (USC) D. Cortina (USC) K. Suemmerer (GSI) para la colaboracion  
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Some nucleosynthesis models developed in the 1940's assumed that nuclide were produced in a primordial process so called fireball, at the beginning of the Universe; those models failed however to explain the experimental observation that stars do not have the same surface composition. Later on, with the help of satellites, the measurement of  $\hat{\text{I}}^3$ -ray coming from the Galaxy confirmed the idea of an ongoing nucleosynthesis scenario still active in stars. The first evidence was the measurement of the  $\hat{\text{I}}^3$  -ray line from the de-excitation of  $^{26}\text{Mg}$  produced by the  $\hat{\text{I}}^2$  -decay of  $^{26}\text{Al}$ . This nucleus has a life time of  $1.05 \times 10^6$  years, much shorter than the age of the Universe, hence the importance of the knowledge of this nucleus and its neighbourhood.

$\hat{\text{I}}^3$  -ray measurements at the Galactic plane showed that  $^{26}\text{Al}$  is mainly produced in massive stars (novae and supernovae). One possible production way is the rp-process.  $^{26}\text{Al}$  has a metastable and a ground state, the first one decays predominantly to the ground state of  $^{26}\text{Mg}$ , nevertheless the ground state decays to the first excited of  $^{26}\text{Mg}$  giving a  $\hat{\text{I}}^3$  -ray of 1.809MeV (the one detected in galactic measurements). The  $\hat{\text{I}}^2$  -decay of  $^{26}\text{Si}$  mainly populates the  $^{26}\text{Al}(\text{g.s.})$ , and the production of the  $^{26}\text{Si}$  comes from the competition of the  $\hat{\text{I}}^2$  -decay  $^{25}\text{Al}$  and the reaction  $^{25}\text{Al}(\text{p},\gamma)^{26}\text{Si}$  that is destructed in the  $^{26}\text{Si}(\text{p},\gamma)^{27}\text{P}$ . Within this scenario the  $^{26}\text{Si}(\text{p},\gamma)^{27}\text{P}$  reaction appears as important, first because it is in the rp-path and secondly because it influences the generation of  $^{26}\text{Al}$ .

The direct study of the reaction  $^{26}\text{Si}(\text{p}, \hat{\text{I}}^3)^{27}\text{P}$  at astrophysical energies is extremely challenging due to the low intensity associated to low energy radioactive beams and low cross sections involved. Coulomb dissociation studies of the inverse kinematics reaction  $^{27}\text{P}(\hat{\text{I}}^3, \text{p})^{26}\text{Si}$  have instead been proposed. A  $^{27}\text{P}$  beam impinges on thick Pb target. The  $^{27}\text{P}$  is then excited via the absorption of a virtual photon to a particle unbound state which decays into  $\text{p}+^{26}\text{Si}$ . This inverse reaction profits of a much larger cross section. The experiment was performed using the ALADIN-LAND setup at GSI with a  $^{36}\text{Ar}$  primary beam at 500MeV. A secondary beam of  $^{27}\text{P}$  was produced by projectile nuclear fragmentation at the FRS. The ALADIN-LAND setup allows to measure in full kinematics. After the Coulomb dissociation of  $^{27}\text{P}$  under the effect of the thick Pb target, both outgoing fragments, protons and  $^{26}\text{Si}$  enter in the large acceptance magnet (ALADIN) and are deflected differently according to their associated rigidities. A set of detectors located after the magnet allow to track and identify protons and fragments in an event by event basis. Preliminary results of the associated invariant mass spectrum will be presented.